



ABSTRACT:

- Uncertainty about the relationship between aeolian saltation flux and wind shear stress hinders understanding of earth and planetary geomorphology and atmospheric dust generation.
- Here, we investigate the saltation flux law based on data from three field campaigns yielding comprehensive data of unprecedented scope.
- Our observations show that mean saltation layer height remains constant with changes in shear velocity, indicating constant mean particle speeds.
- Based on this, we predict a linear relationship between saltation flux and shear stress scaled by the typical duration of saltation hops.
- Direct stress-flux comparison strongly supports inference of flux law with saltation flux increasing linearly with excess shear stress.
- We expect a linear relationship to hold in other environments (e.g. Mars) with large particle-fluid density ratios.

SALTATION FLUX LAW DEBATE (LINEAR VS. NONLINEAR 3/2):

- Q (g m⁻¹ s⁻¹) saltation mass flux
- au (Pa) shear stress
- $\tau = \rho_f u_*^2$, where ρ_f is air density and u_* is shear velocity $Q = \Phi V$ (Particle concentration * Particle speed)
- Mean particle concentration Φ scales linearly with excess stress au_{ex} , $au_{ex} = au - au_{it}$ (e.g., Ho et al., 2011)
- Mean **particle speed** *V* is controversial:
 - Most existing flux laws (e.g., <u>Bagnold</u>, <u>1941</u>; Owen, <u>1964</u>) **assume** $V \sim u_*$ and therefore predict $Q \sim \tau^{3/2} (Q \sim u_*^3)$
 - Recent theory (e.g. Ungar and Haff, 1987) and experiments (e.g. Ho et al., 2011) alternatively support V constant with u_* and therefore predict $Q \sim \tau (Q \sim u_*^2)$
- Further, V is related to saltation layer height z_q as: $V \sim \sqrt{z_q}$ (Owen, 1964)
- Therefore, determining z_a versus u_* should resolve Q vs. τ scaling

FIELD CAMPAIC	<u>ins</u>		
	Jericoacoara (Ceara, Brazil)	Rancho Guadalupe (California, USA)	Oceano (California, USA
# saltation days	3	2	12
surface d_{50}	0.55 ± 0.04	0.53 ± 0.03	0.40 ± 0.07
 Saltation flux, <i>C</i> calculated from exponential fit for 30 minute avg for profile of Wenglors (<i>z</i> ≈ 247 cm), shown right, calibrated from ~hourly B sand trap collections. Shear stress, <i>τ</i>, computed as Reynolds stress (<i>τ</i> = −ρ_f <i>u'w'</i>) over 30-minute windows from sonic anemome mounted at <i>z</i> ≈ 0.5 m. 	2, 1, 2- on 5NE SNE SNE (a) Jer Guadalupe the side.	icoacoara, looking upwind, (c) Oce (d) Close-up of Wenglor a looking downwind	Image: constraint of the second se

Linear scaling of wind-driven sand flux with shear stress

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INTERPRETING THE LINEAR FLUX LAW

Our observations provide the first strong field-based evidence for a linear (as opposed to 3/2) flux law.

- *Past field work:* Insufficient data to resolve flux law (e.g., Sherman and Li, 2012)
- Past experiments: Insufficient development of saturated flux (e.g., Rasmussen et al., 2015) or suppression of
- saltator hops (e.g., Li and McKenna Neuman, 2012). Constant saltation height / particle velocity arises from

dominance of splash entrainment mechanism which requires constant near-surface particle speeds at steady-state

- Therefore, expect constant saltator heights and linear flux law for all cases with particle entrainment dominated by splash:
 - In aeolian case, splash entrainment dominates and V is constant and $Q \sim \tau (Q^2 u_*^2)$.
 - However, In fluvial case (e.g., Lajeunesse et al., 2010),
 - fluid entrainment and $V \sim u_*$ and $Q \sim \tau^{3/2}$ ($Q \sim u_*^3$).
 - *Hypothesis*: Contrast is set by particle-fluid density ratio, $s=
 ho_p/
 ho_f$

sion to planetally surfaces (e.g., Pantz and Duran, 2010	sion	to	planetary	surfaces	(e.g.,	Pahtz	and	Duran,	2016
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nent	Particle-fluid density ratio: $s = \rho_p / \rho_f$	Flux law type
<i>r</i> ial	2.65	Nonlinear 3/2
	40	?
	190	?
olian	2000	Linear
	2.5×10^{5}	Linear predicted
	10 ⁷	Linear predicted
	10 ⁷	Linear predicted
	10 ¹²	Linear predicted
	e.g., 10 ¹²	Linear predicted

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